

# Examining Interchange at the Outer Planets using JERICHO: a Kinetic-Ion, Fluid-Electron Hybrid Plasma Model

13<sup>th</sup> July 2022 Josh A. Wiggs<sup>1</sup>, C. S. Arridge<sup>1</sup> & C. Paty<sup>2</sup>

<sup>1</sup>Department of Physics, Lancaster University, UK <sup>2</sup>Department of Earth Sciences, University of Oregon, USA







### **Radial Interchange at the Outer Planets**





Have plasma sources in inner magnetospheres of Jupiter and Saturn

Plasma is transported into middle and outer magnetosphere by RI instability

Liu et al, 2010

What's happening inside here? Region is too small to probe with current state-ofthe-art

$$\underline{F}_{cf}\cdot\nabla\int\left(\frac{\rho}{B}\right)ds\geq 0$$

Southwood & Kivelson 1987, Achilleos+ 2015

#### JERICHO – Model Topology





#### 4

#### JERICHO – Model Topology







#### JERICHO – Kinetic Ion, Fluid Electron





-5

#### What controls size and propagation of instabilities

- Initial conditions set for Saturn (Wilson+ 2017, Persoon+ 2020)
- Left & right boundaries are periodic, plasma source placed at bottom (7.5R<sub>s</sub>), outflow allowed at top (9.5R<sub>s</sub>) (*Azari+ 2018*)
- Plasma source pulsed over course of simulation run, enhancements in plasma density induces interchange instability



Space &





#### Radial-Interchange Instability – Model Run

Space & Lancaster 🏁 Planetary Physics



7

#### Radial-Interchange Instability – Initial Analysis



On order 10 instabilities form on front of plasma injection

Space &

Planetary Physics

Lancaster University

- No gravity included and no velocity shear along interface of injection
- Narrow channels of tenuous plasma form between dense fingers

Magnetic Field?



- Varied strength of initial magnetic dipolar field
- Both length and temporal scale for development of instabilities remains constant
- Stronger fields act to retard outwards flow of injected plasma



### Plasma Source Density?





- Varied density of plasma injected from source at boundary
- Both length and temporal scale for development of instabilities remains constant
- Denser injections spread to occupy larger domain spaces

### Summary



- We have developed a 2.5D ion-kinetic, fluid-electron hybrid plasma model in c++ [*in prep, Wiggs & Arridge*]
- Plasma source placed at bottom injects high density region on the leading and trailing edge of which  ${\sim}10$  instabilities form ( ${\sim}0.1R_{\rm S})$  spontaneously form
- Varying both magnetic field strength & injected plasma density does not significantly alter spatial or temporal scales
- Development ongoing Parallelise to increase simulation size (RAM bound), polar-cylindrical coordinates, ionosphere, Jovian ICs





JoshWigg



#### Backup

#### Parameter Survey – Resolution Change

Space & Lancaster 🏁 Planetary Physics University



## What controls size and propagation of instabilities

- Initial conditions set for Saturn (*Wilson+ 2017, Persoon+ 2020*) between 7.5R<sub>s</sub>-9.5R<sub>s</sub> (*Azari+ 2018*)
- Periodic boundaries on left-right boundaries, plasma source places at bottom, outflow set at top
- Plasma source pulsed over course of simulation run, enhancements in plasma density induces interchange instability



Wilson+ 2017

